

The Concatenator: A Bayesian Approach To Concatenative Mosaicing

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Motivation

- Ben found Chris online by his open source implementation of Let It Bee^[4]
- Ben^[6] and others^[5,7] were making awesome music with Driedgers' technique^[1], but it was too slow. We wanted a similar effect on **large corpora** in **real time**
- Music producers have so many sample packs on their computer that they haven't even listened to all of the sounds. We want a system that can efficiently pick out sounds for them to match with a target, preserving rhythm and pitch

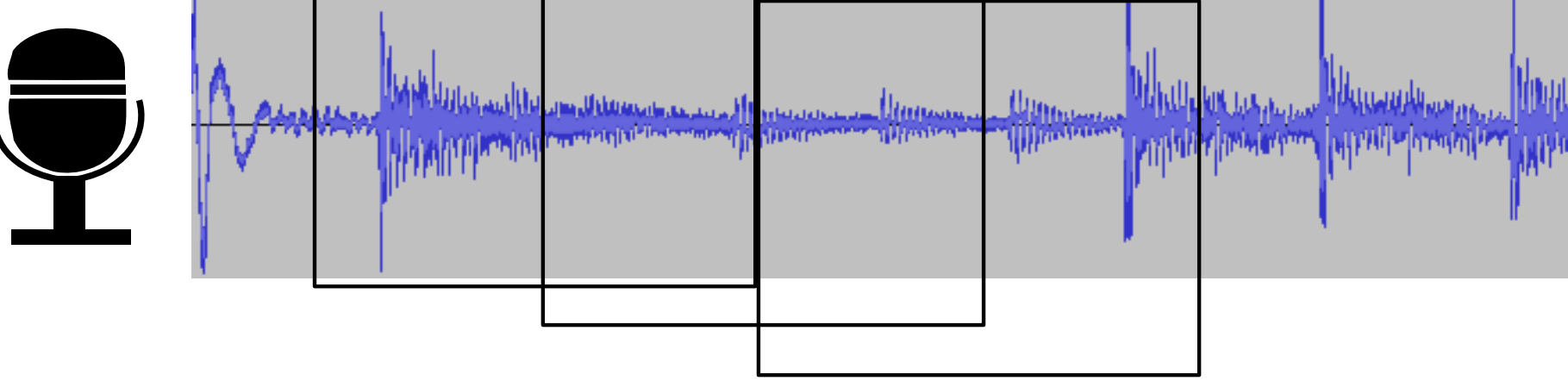
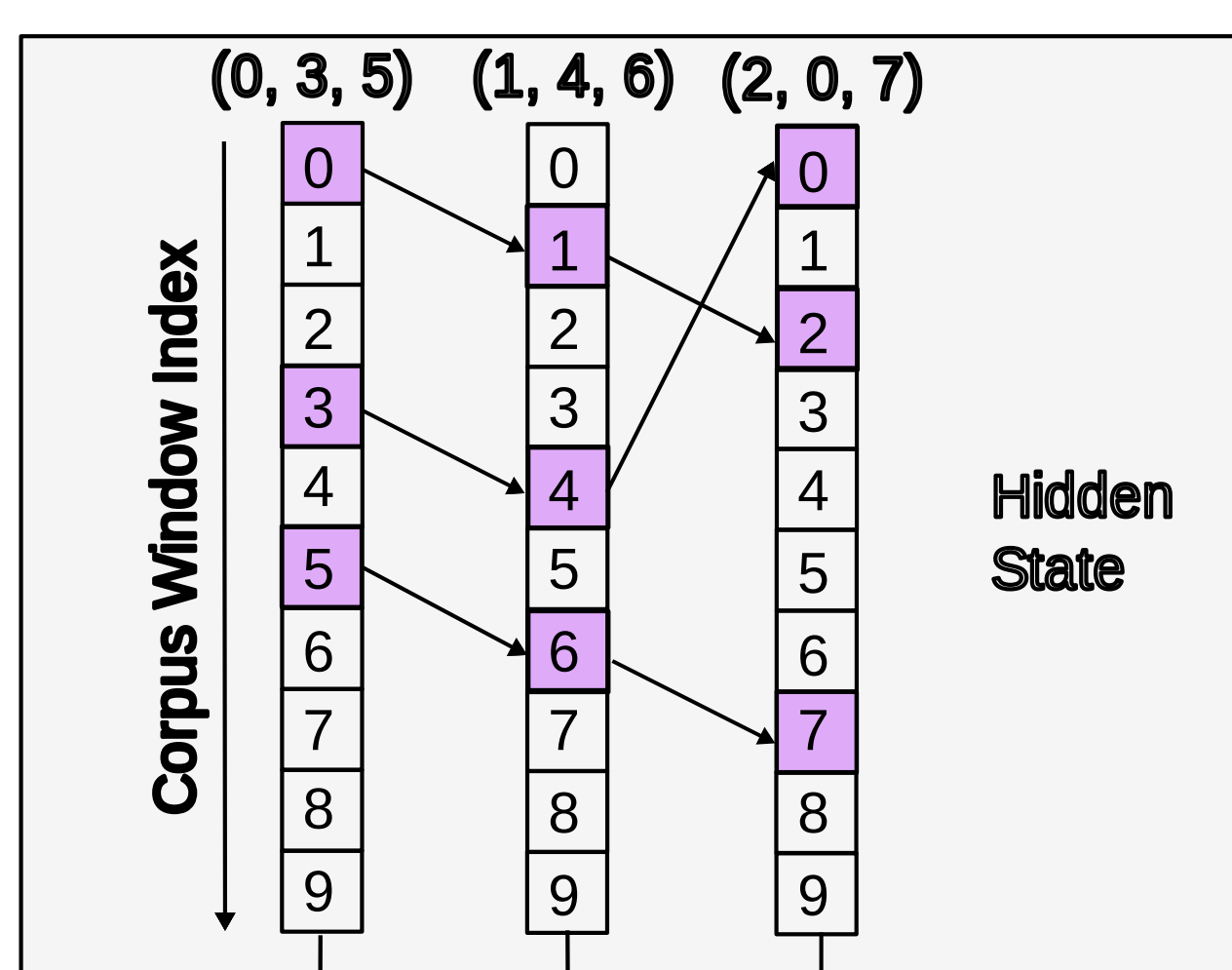
State Space

- A state s_t at time t represents a **sparse** selection out of N possible corpus windows as a p -length vector of indices

$$\vec{s}_t[k] \in \{0, 1, \dots, N-1\}, k = 0, 1, \dots, p-1$$

- Given associated weights \mathbf{h} , and a spectrogram \mathbf{W} for the corpus, the spectral approximation of the target at frequency bin m is

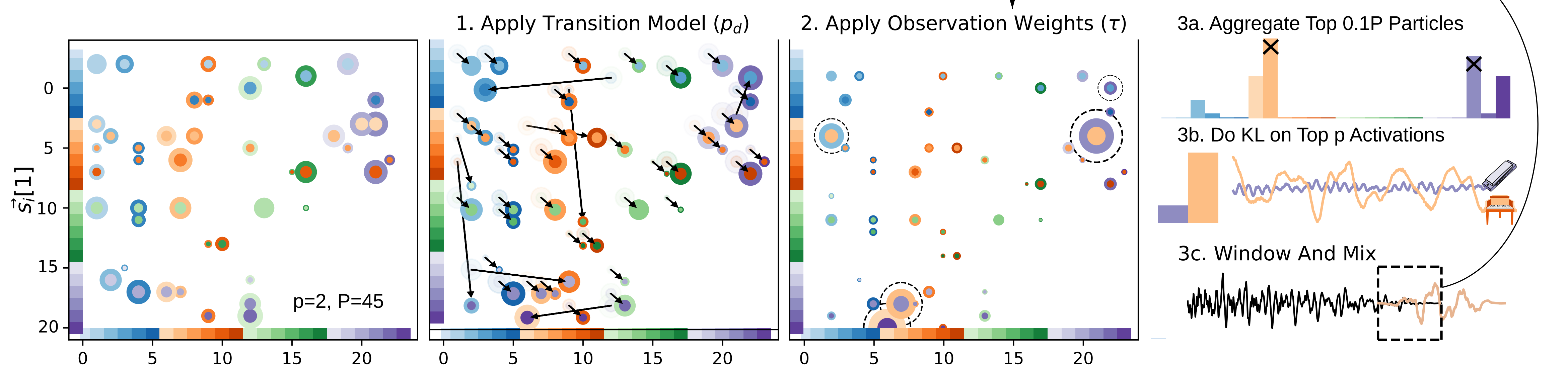
$$\vec{\Lambda}_t[m] = \sum_{k=0}^{p-1} \vec{h}_t[k] W_{m, \vec{s}_t[k]}$$



Particle Filter Pipeline for Concatenative Mosaicing

Beginning: Randomly throw a bunch of "particles" (darts) into the corpus. Then repeat the following steps as time progresses

- With high probability p_d , slide particles forward in time in the corpus. With small probability $(1-p_d)$, randomly jump to new locations in the corpus to adapt to new sounds coming in.
- Give higher weight to particles that fit the target better, according to KL divergence of the best possible fit *using only the activations from that particle*
- Resample particles periodically according to weight to replicate ones with a good fit ("survival of the fittest")
- At every hop length time increment, have the top 10% particles vote to determine the final corpus activations, and solve one final NMF problem to mix in audio



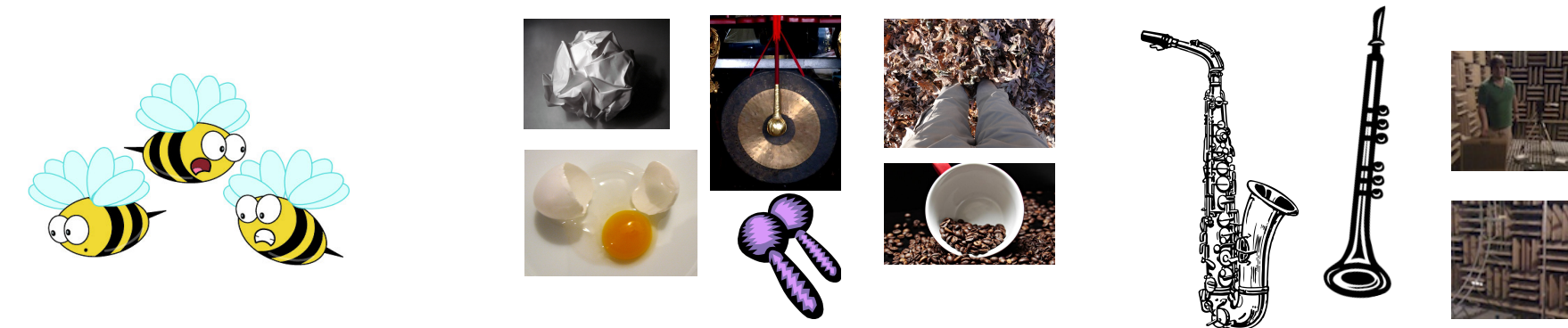
Driedger^[1] NMF Time Complexity: $O(LMNT)$: *Linear* dependence on corpus size

The Concatenator Time Complexity $O(LMPpT)$; **Does not scale with corpus size!** (We also use smaller L)

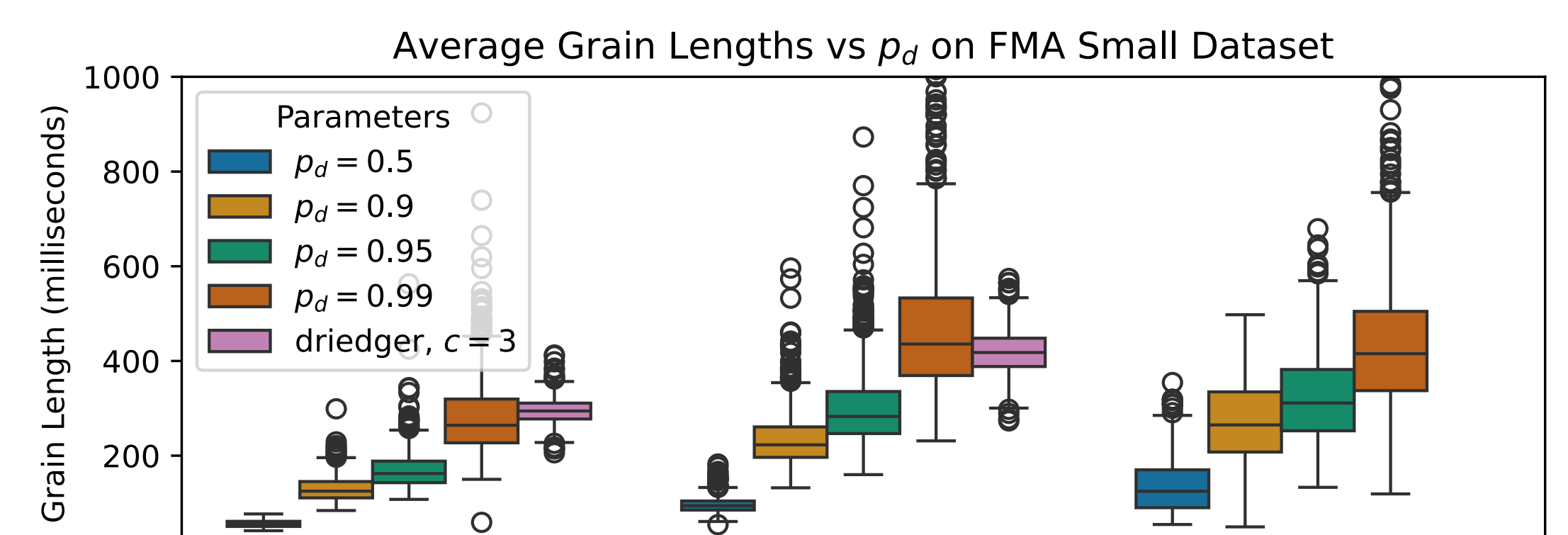
N: Number of corpus windows
M: Number of spectrogram bins
P: Number of particles
p: Activations per window
L: NMF Iterations (we can use fewer than Driedger)
T: Number of windows in target

Quantitative Evaluation: Free Music Archive (FMA) Dataset

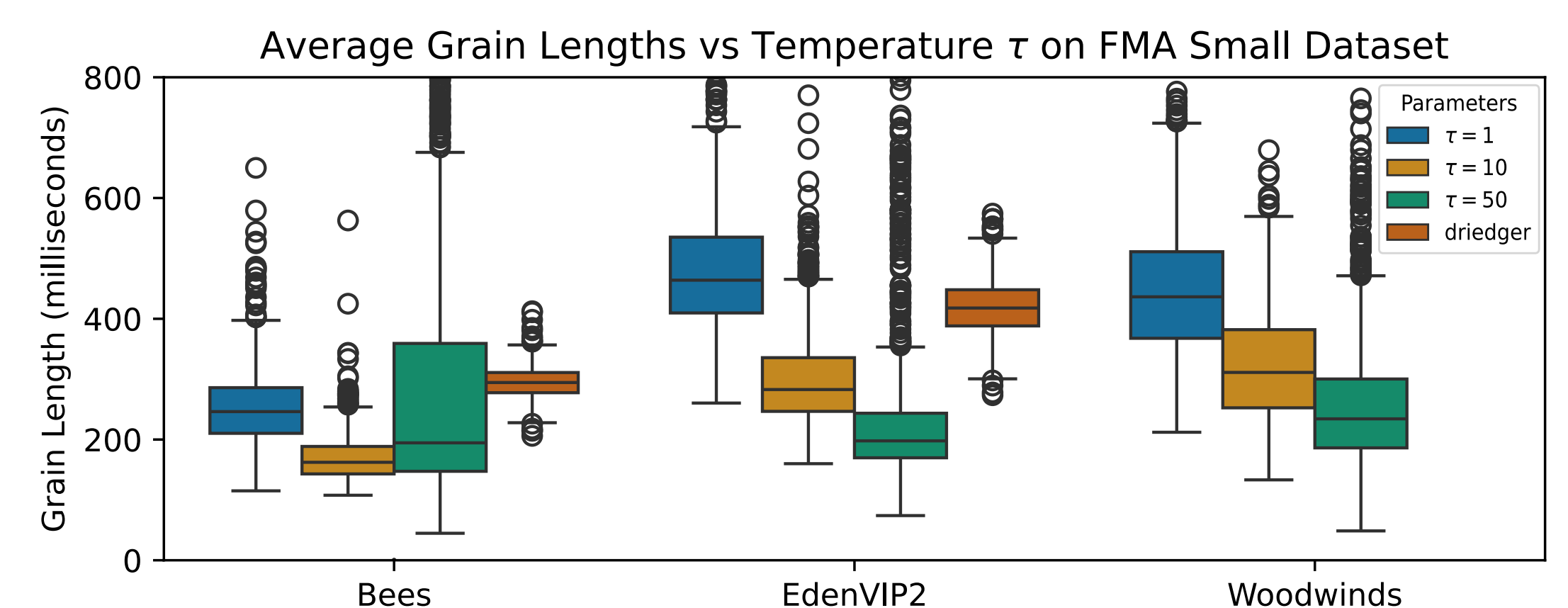
- Randomly subsampled 1000 30 second clips from the Free Music Archive (FMA)-small dataset, each of which we used as a target for 3 different corpora:



Small: Driedger Bees^[14] (~1 minute)
Medium: Eden VIP2^[6] (~10 minutes)
Large: Ulowa Woodwinds^[12] (~1.6 hours)



- Increasing p_d **increases** the average grain length since windows are less likely to jump at each timestep.



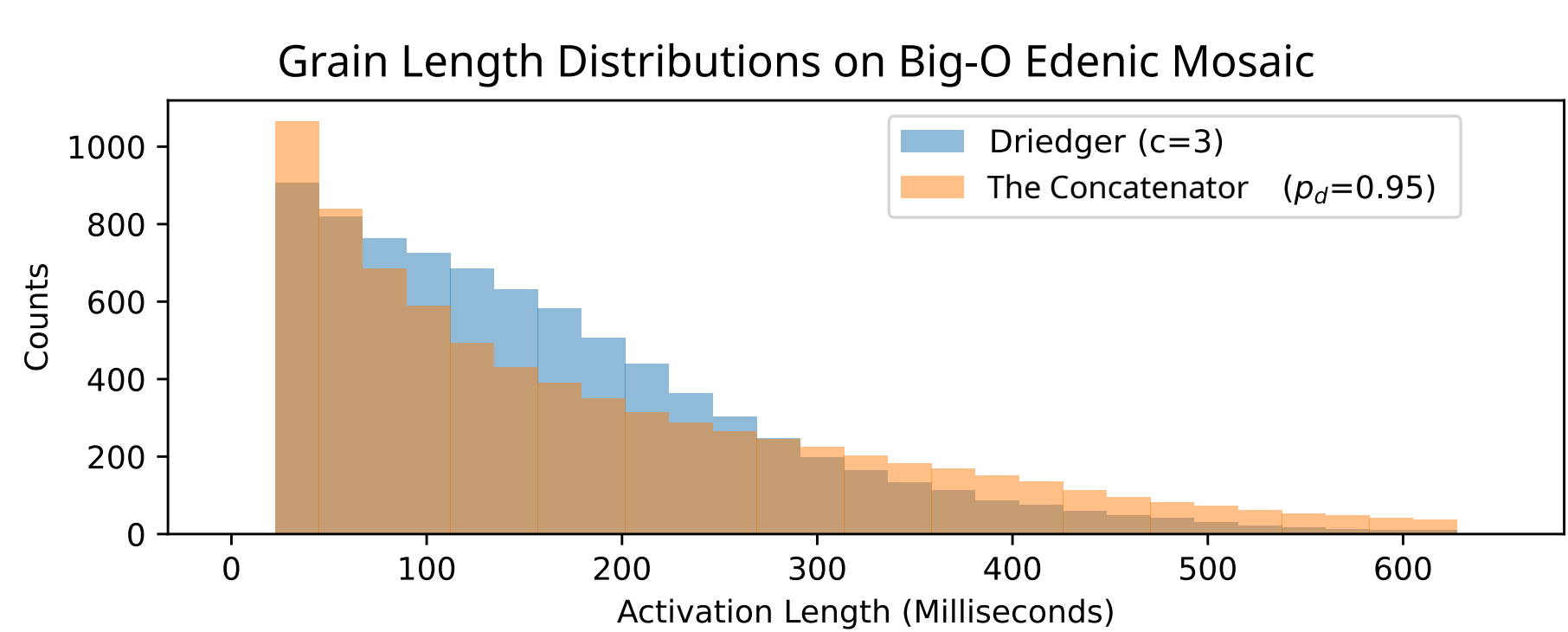
- Increasing temperature τ **decreases** the average grain length since this prioritizes the observation probability at each timestep.

Transition Model

- Markov assumption:** This state only depends on the last state
- Move to the next window in the corpus in time order with probability p_d
- Different activations transition independently; **factorial hidden markov model**^[13]

$$P_T(\vec{s}_t = \vec{b} | \vec{s}_{t-1} = \vec{a}) = \prod_{k=0}^{p-1} \begin{cases} p_d & \vec{b}[k] = \vec{a}[k] + 1 \\ \frac{1-p_d}{N-1} & \text{otherwise} \end{cases}$$

- In the absence of observation corrections, leads to a geometric distribution of grain lengths



Observation Model

- Cost for the i^{th} particle based on the target spectrogram \mathbf{v}_t is:

$$D_i(\vec{v}_t | \vec{\Lambda}_i) = \left(\sum \vec{v}_t \odot \log \left(\frac{\vec{v}_t}{\vec{\Lambda}_i} \right) - \vec{v}_t + \vec{\Lambda}_i \right) + \frac{\|\alpha \odot \vec{h}_i\|_2^2}{2}$$

- Important to apply L2 regularization for near silence with a factor α
- To find the activations \mathbf{h}_i , minimizing D_i for each particle, we perform $L=10$ iterations of the regularized KL update equation for NMF^[10], **using only the activations \mathbf{s}_i for this particle**

$$\vec{h}_i^\ell[k] \leftarrow \vec{h}_i^{\ell-1}[k] \left(\frac{\sum_m (W_{m, \vec{s}_i[k]})(\vec{v}_t[m]) / (\vec{\Lambda}_i^{\ell-1}[m])}{\sum_m W_{m, \vec{s}_i[k]} + \alpha[k] \vec{h}_i^{\ell-1}[k]} \right)$$

- This is very fast for each particle, and it is embarrassingly parallelizable

- Finally, we convert each divergence to an observation **probability** over all particles using softmax with "temperature" parameter τ , where higher τ means we promote a better fit to the target

$$p_O[i] = \frac{e^{-\tau D_i}}{\sum_j e^{-\tau D_j}}$$

Mixing

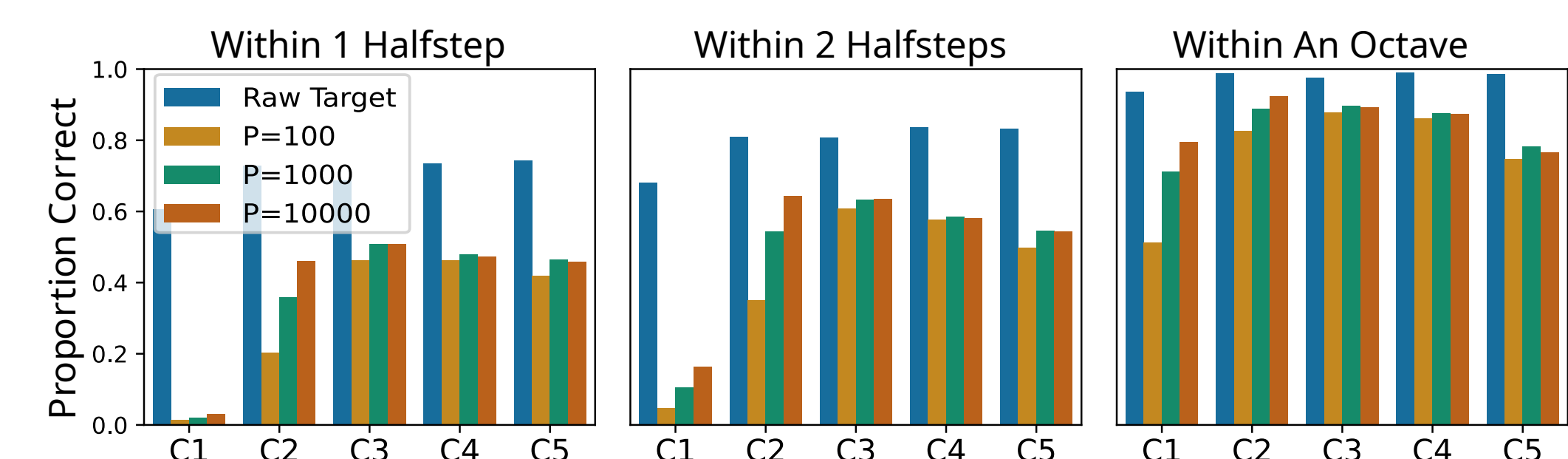
- Audio is mixed together using the same weights learned for spectral fits, using a Hann window (similar to [2] and [8])
- We can choose to zero out repeated activations for some amount of time when the particles vote on the final activations. Driedger is also mindful of this

Probability of Choosing "Good Enough" Particles

$$1 - \left(p_d + (1-p_d) \frac{(N-1-wk)}{N-1} \right)^{(2\delta+1)pP}$$

Allowed time offset from ideal window
Top w similar windows to ideal window

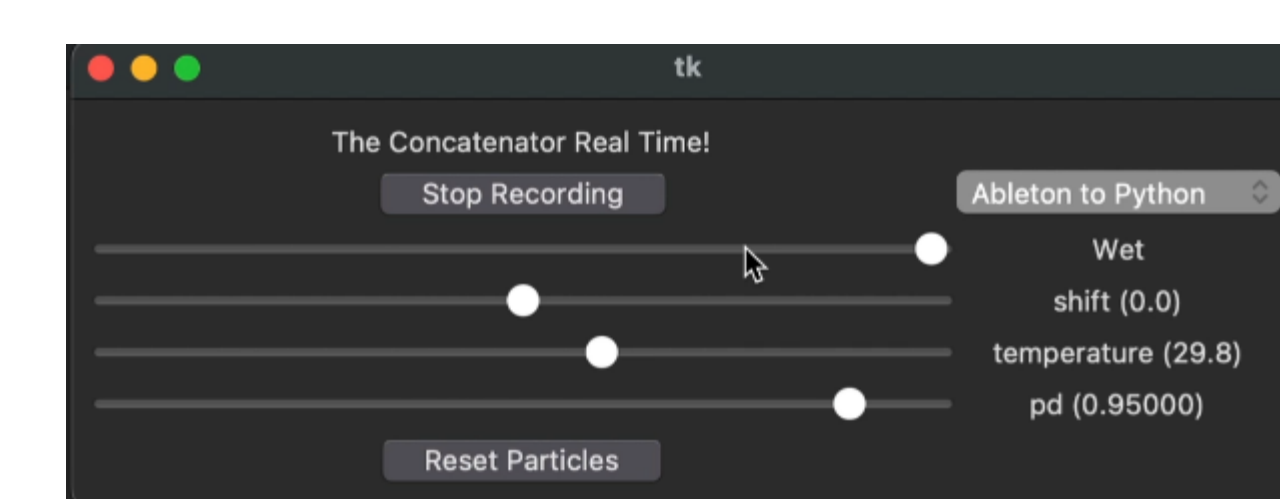
Quantitative Evaluation: Pitch Preservation



- Tested out on stems in Musdb-18hq dataset^[14]
- Good pitch preservation in all but the lower octaves (could be improved with CQT)

Qualitative Results / Open Source Plugin

<https://www.ctralie.com/TheConcatenator>



Built on portaudio^[15] for real time mode

- Huge corpora (e.g. *The Wall* album, hours long sample packs)
- Corpora retaining themselves
- Various target melodies, counterpoint, full mixes, basses, drums, vocals, noise

DataMind Audio Plugin Coming Soon...



- Special thanks to **Robin Leathart** for tireless software development!
- Special thanks to **Jacob Crider** for UI layout and skin design
- Special thanks to **Luca Ayscough** for help optimizing the initial prototype
- Special thanks to **Yashique Chalil** and **Martin Parker** for value feedback during beta testing

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